

**Simultaneous VLBA polarimetric
observations of the $v=\{1,2\}$ $J=1-0$ and
 $v=1$, $J=2-1$ SiO maser emission towards
VY CMa: maser morphology and
pumping**

L. Richter, A. Kembball and J. Jones
MNRAS 436, 1708, 2013

2017-6-14

Observations and data reduction

2003-12-20 and 23: Epoch 1

2007-3-5 and 19: Epoch 2

How to align the SiO maser maps of different transitions

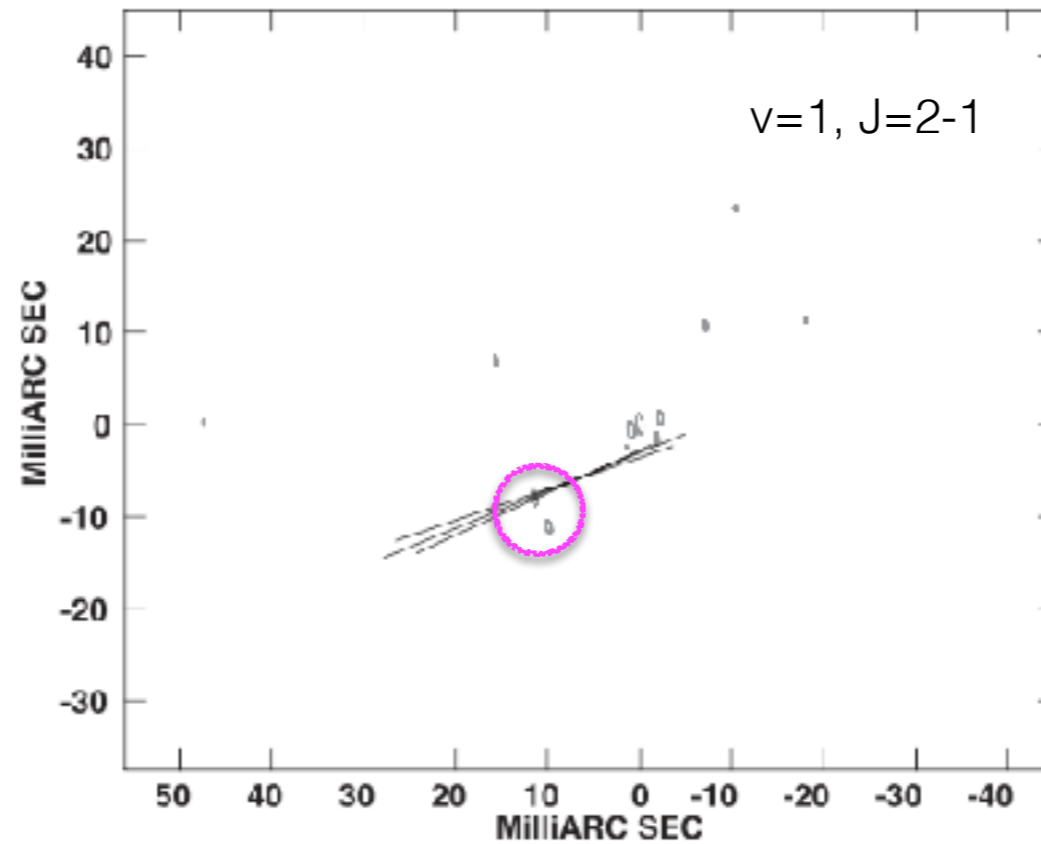
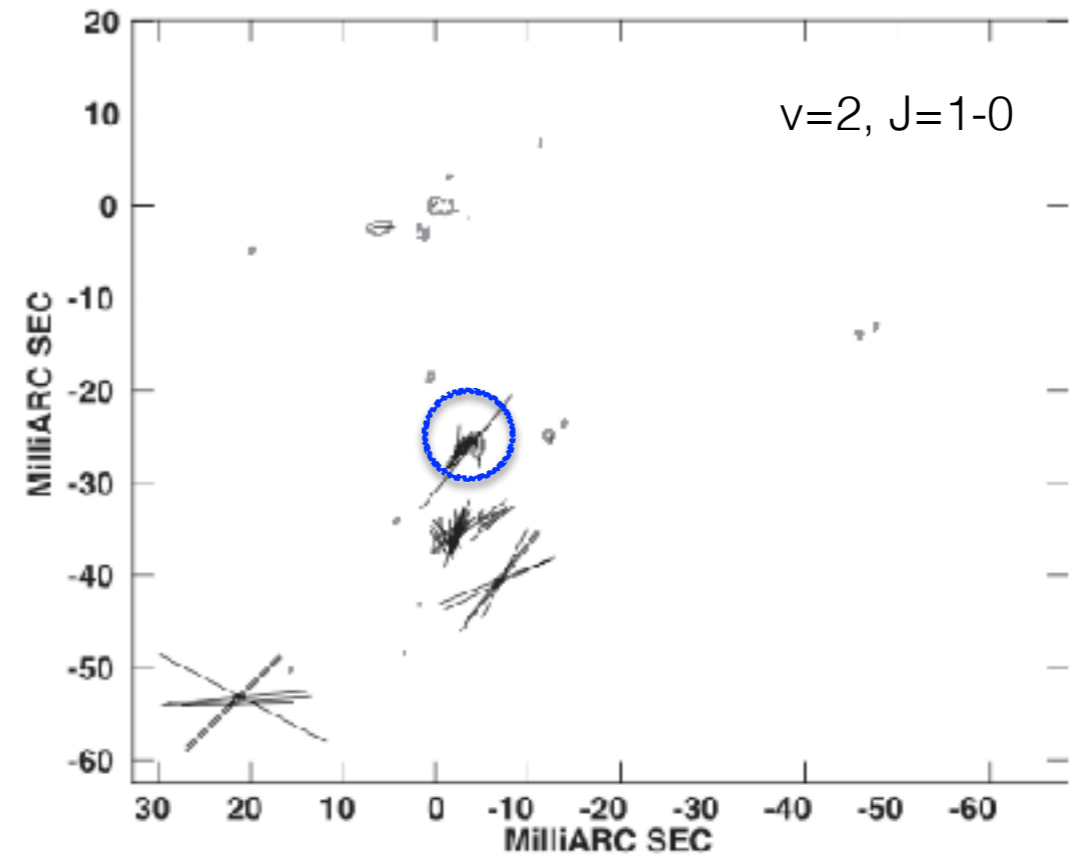
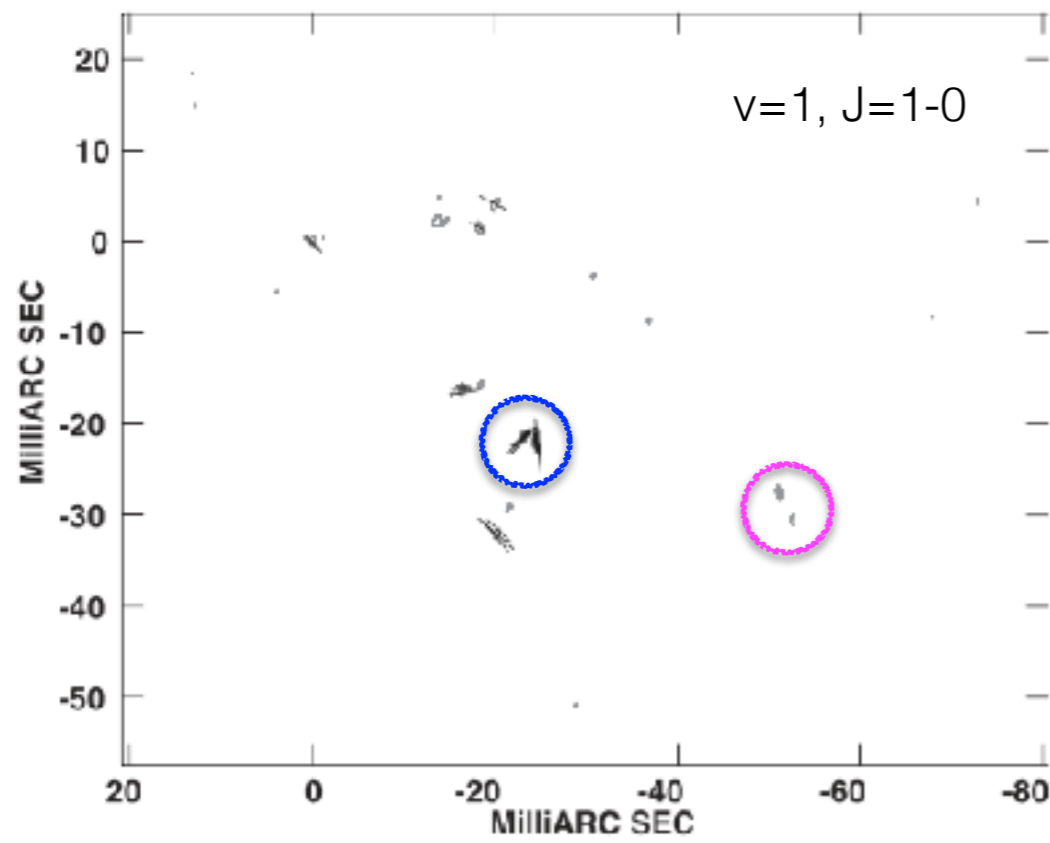
Phase self-calibration: absolute astrometric information lost

Cross-correlation method: uncertainty of map alignment in each epoch < 0.05 mas

Epoch 2: Overlaid transition maps were restored with the same $v=1$ $J=1-0$ beam size. \rightarrow aids component-level interpretation

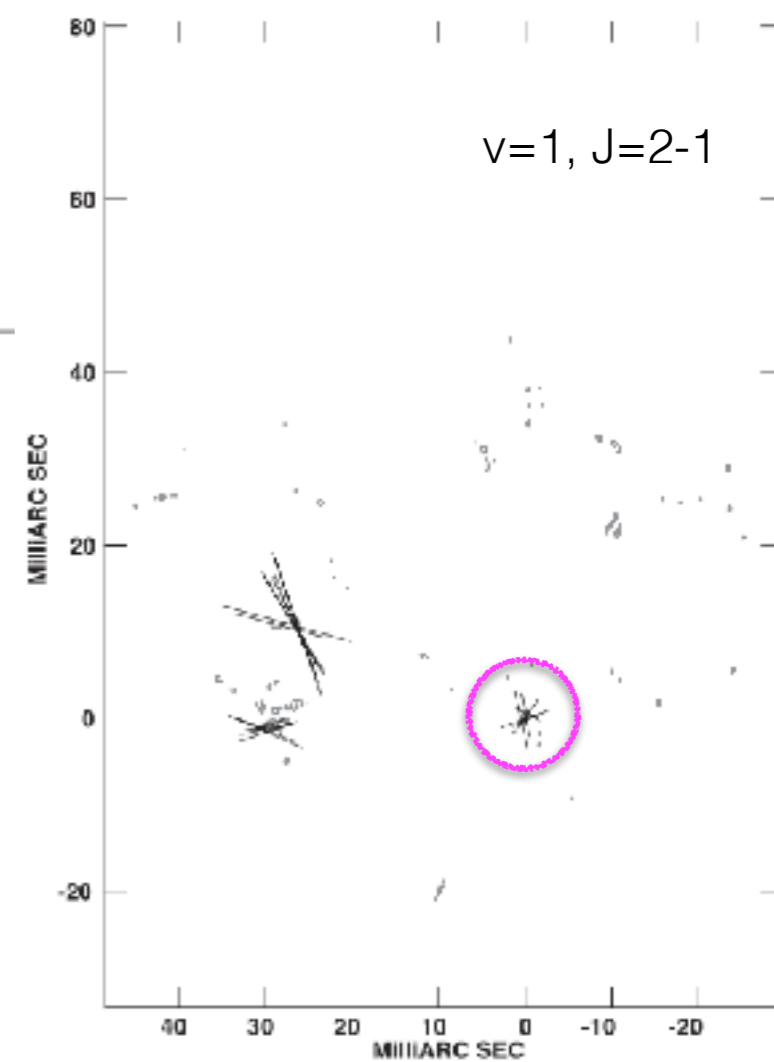
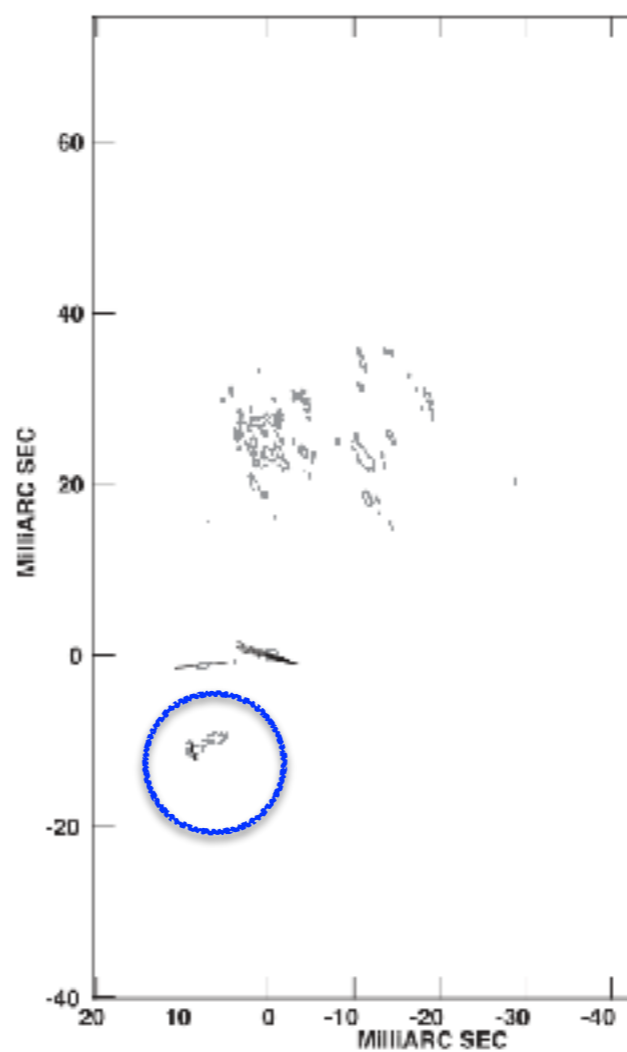
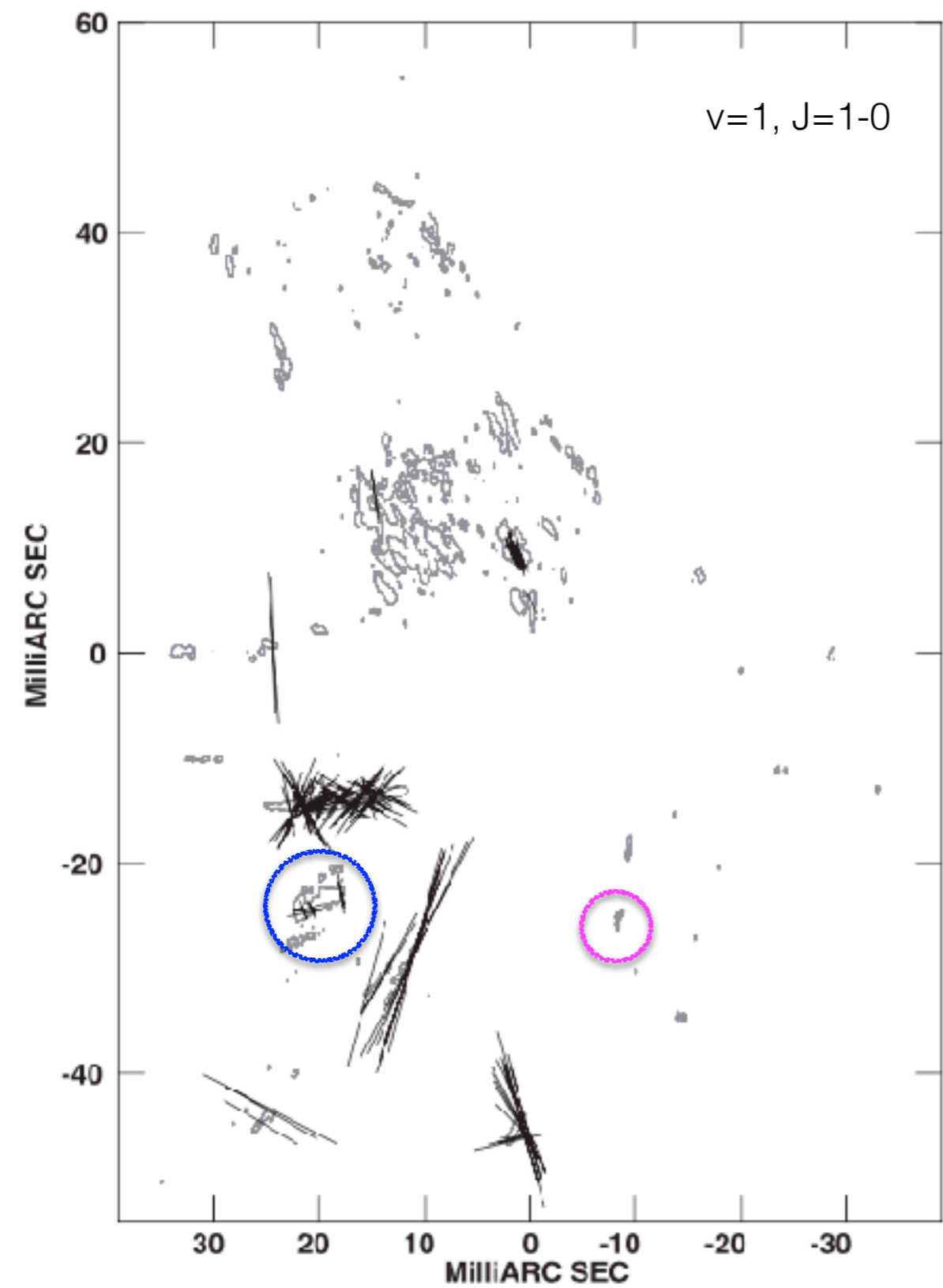
Epoch 1: Overlaid transition map was restored with the intrinsic beam size of each deconvolve map.

Results and discussion



Epoch 1

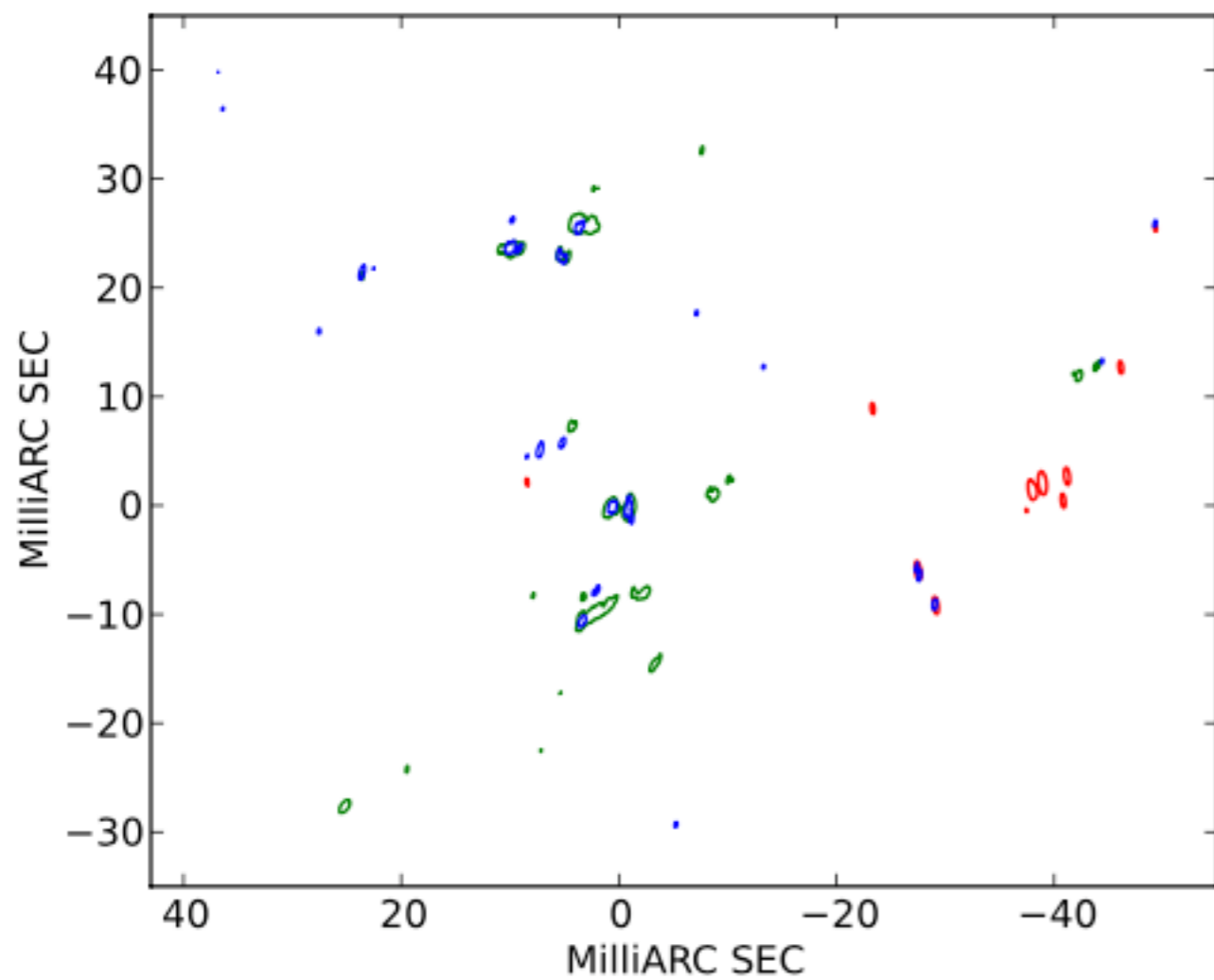
Results and discussion



Epoch 2

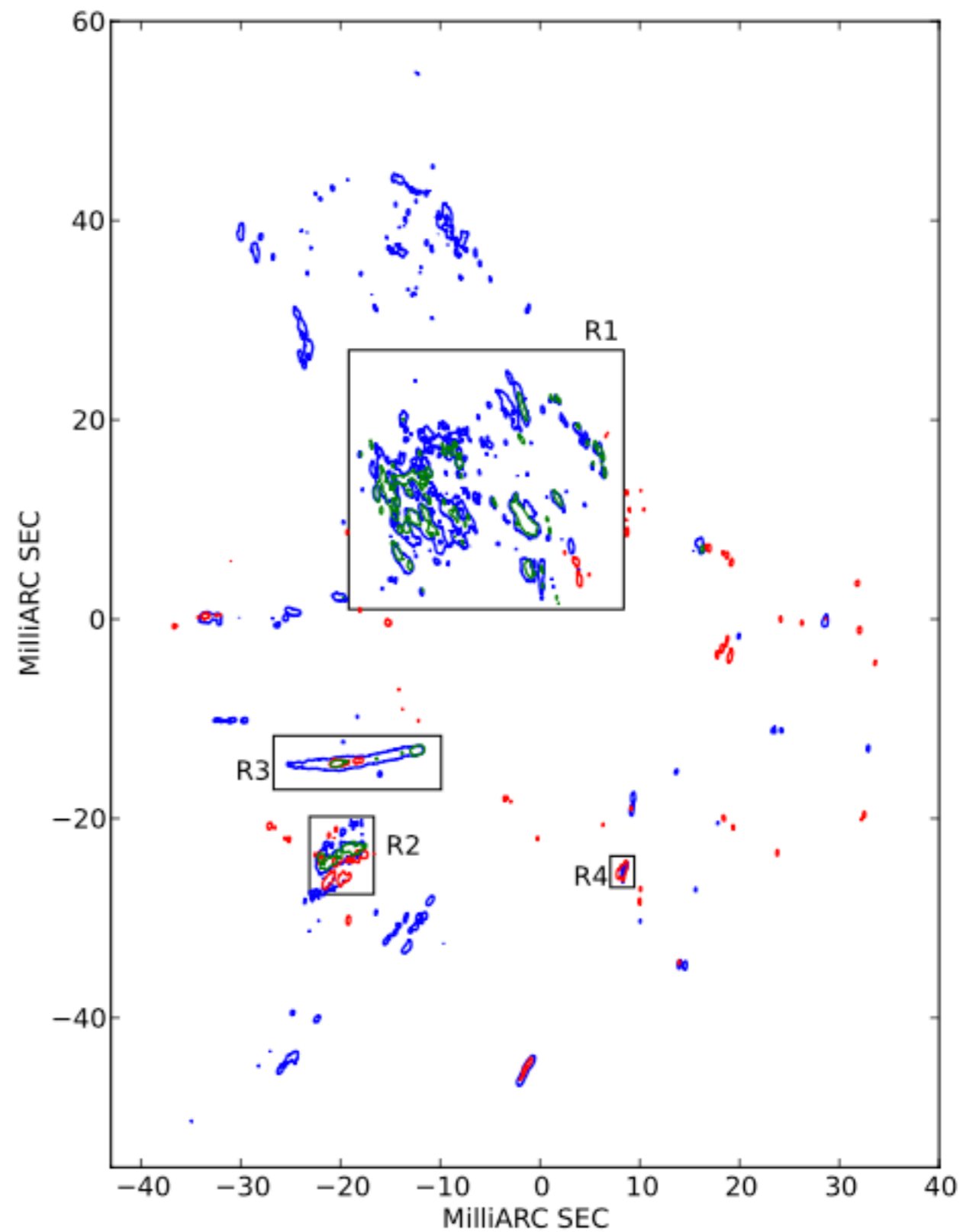
Results and discussion

Epoch 1

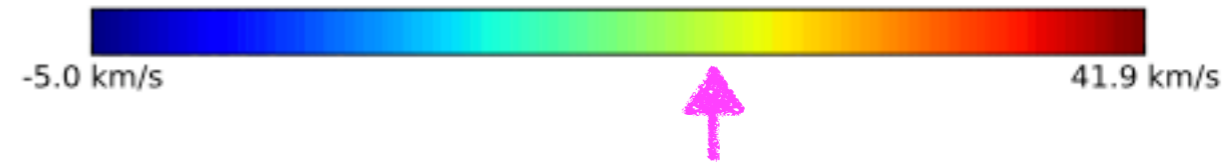
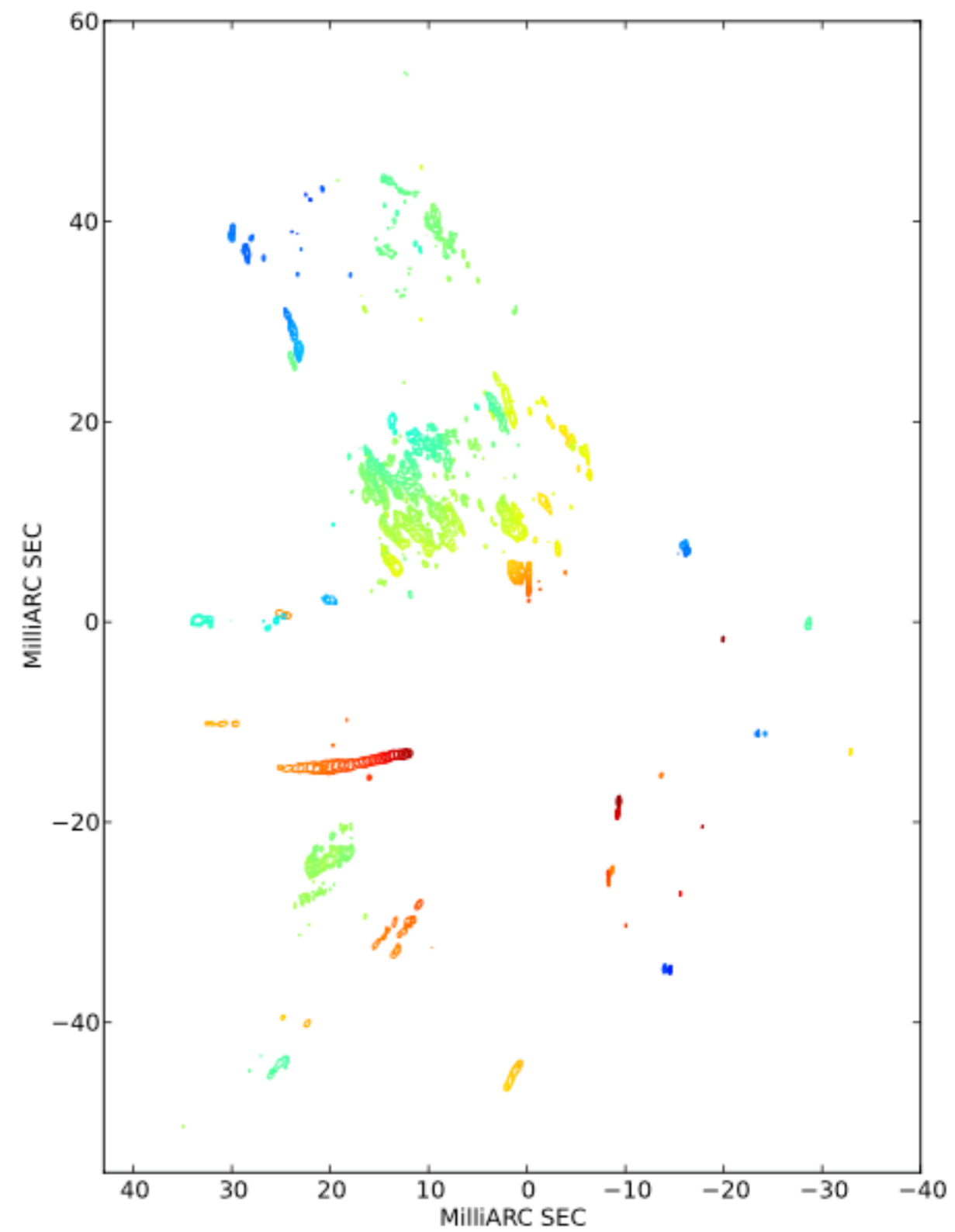
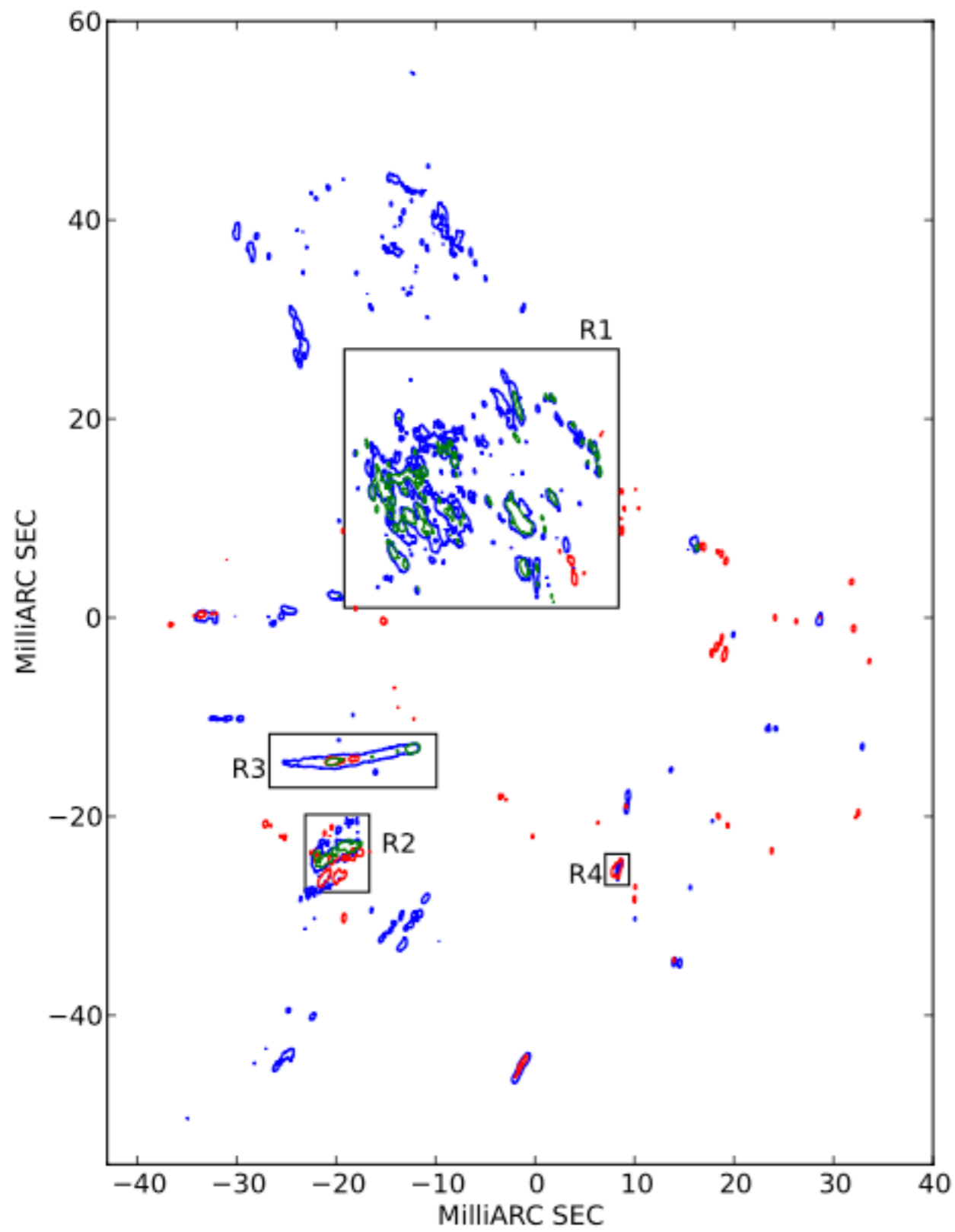


Blue: $v=1, J=1-0$
Green: $v=2, J=1-0$
Red: $v=1, J=2-1$

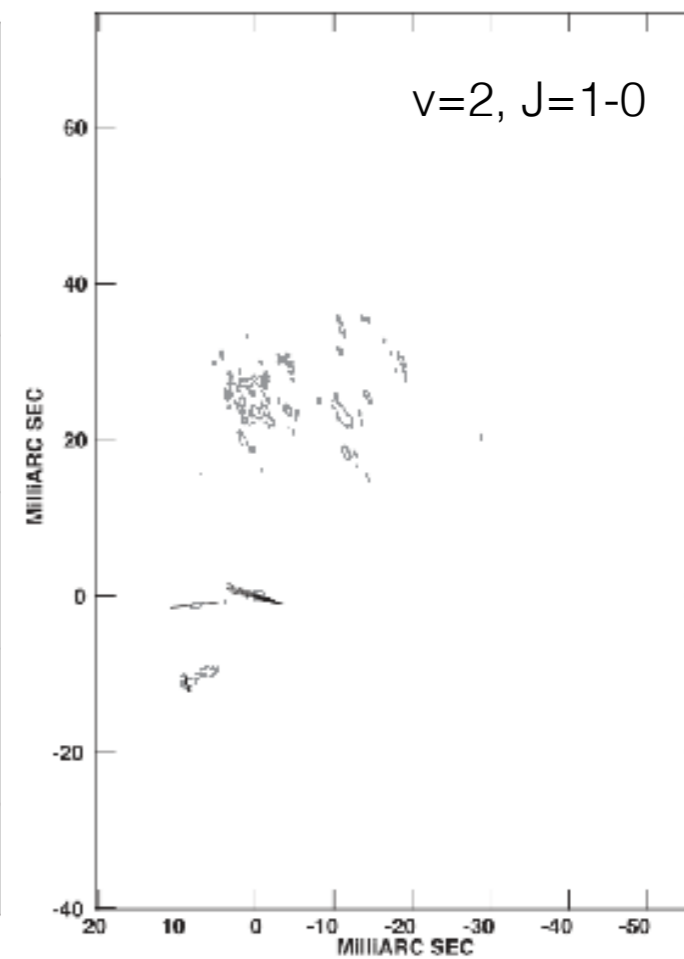
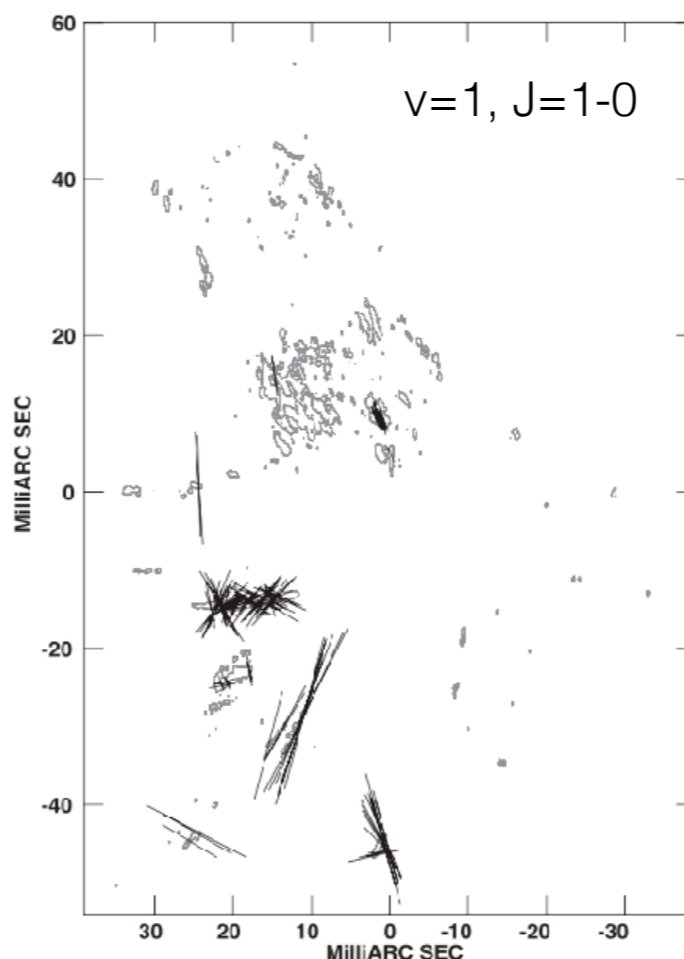
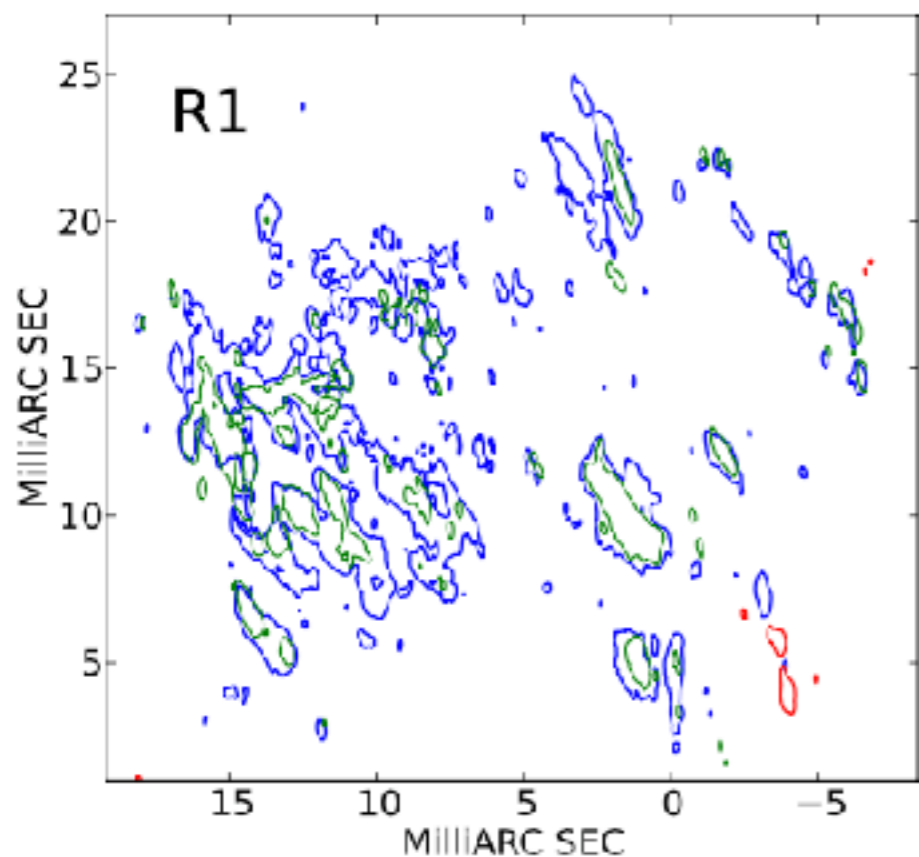
Epoch 2



Results and discussion

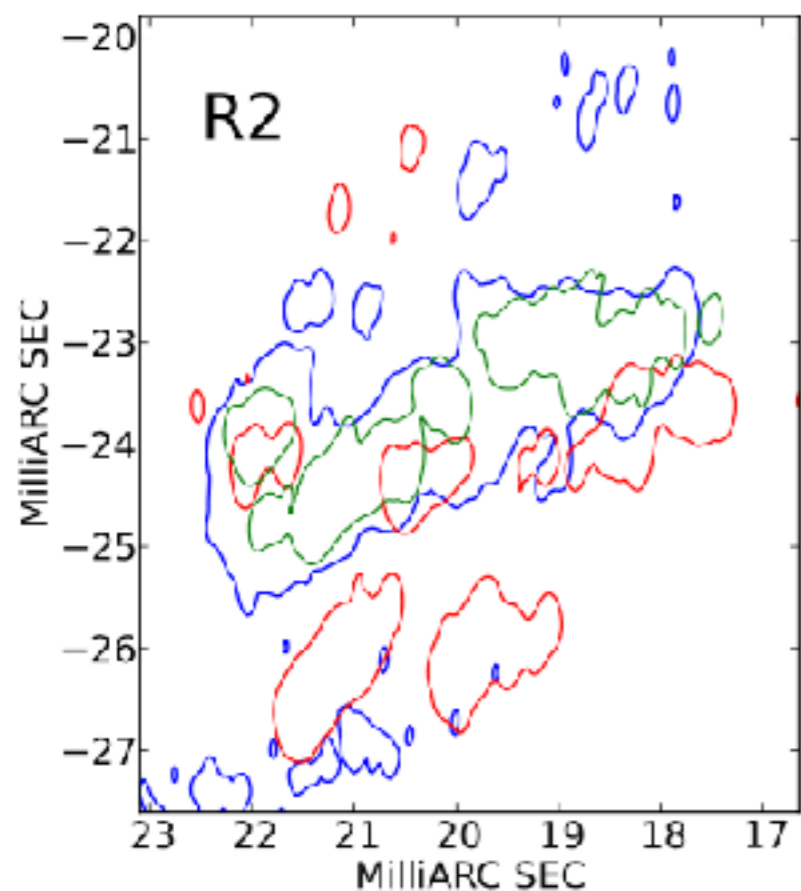


Results and discussion



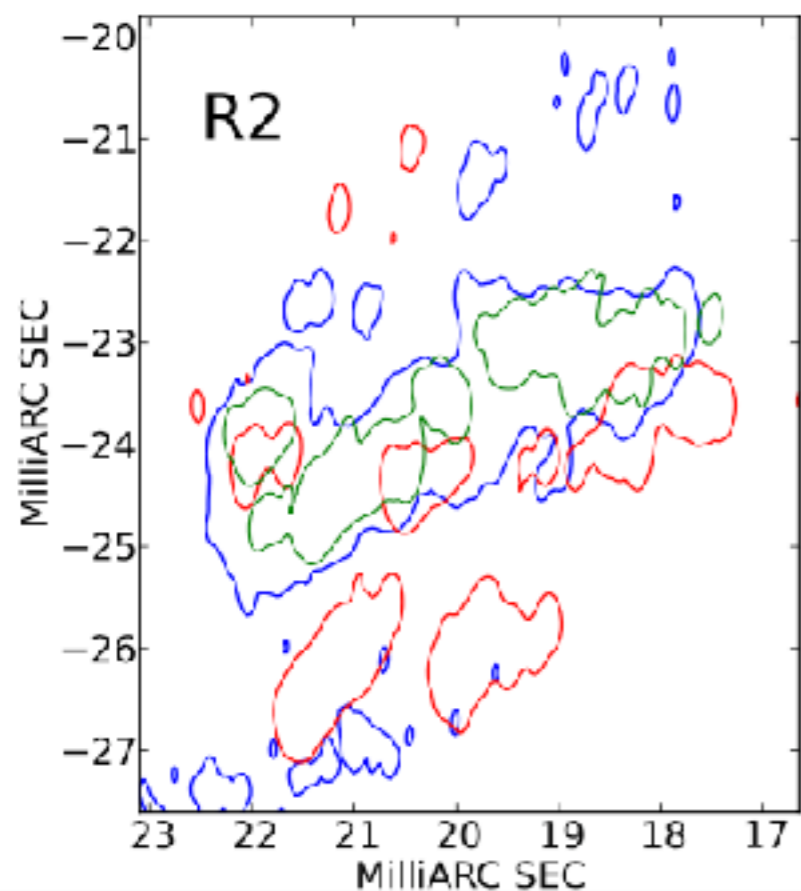
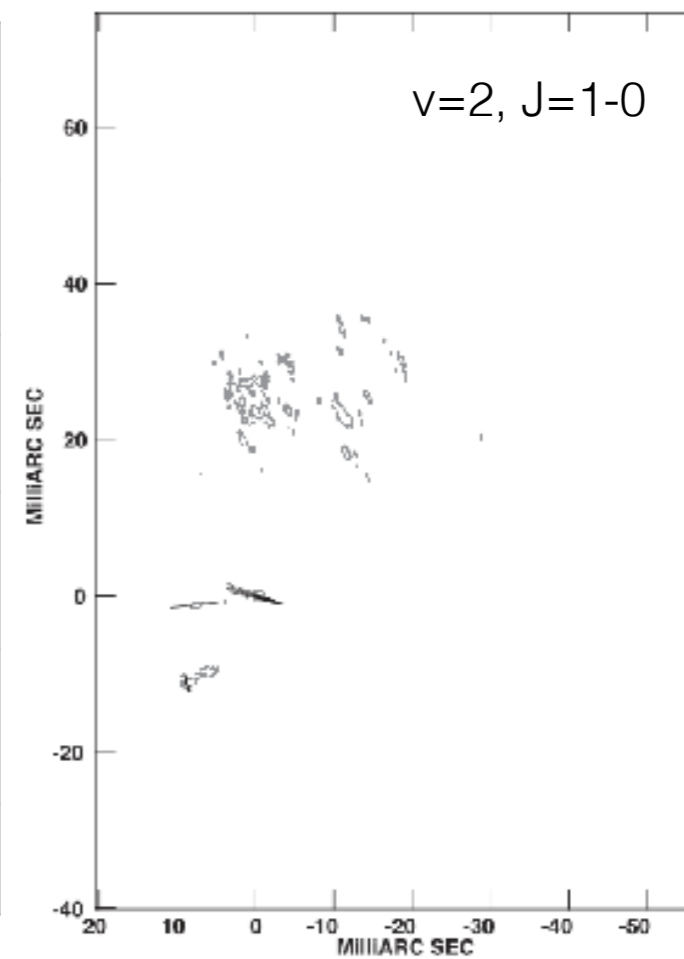
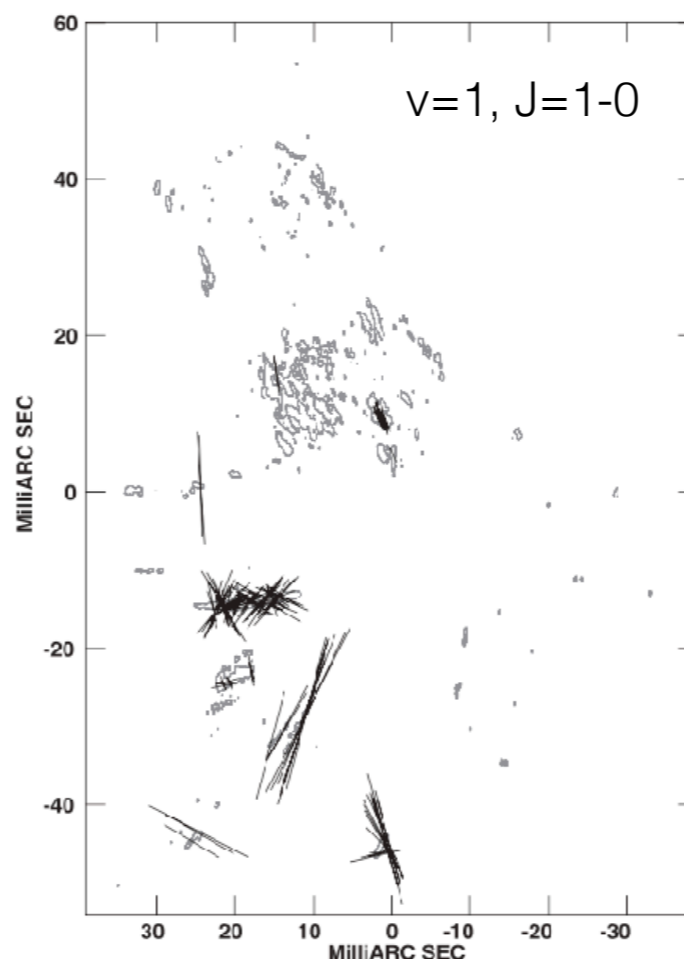
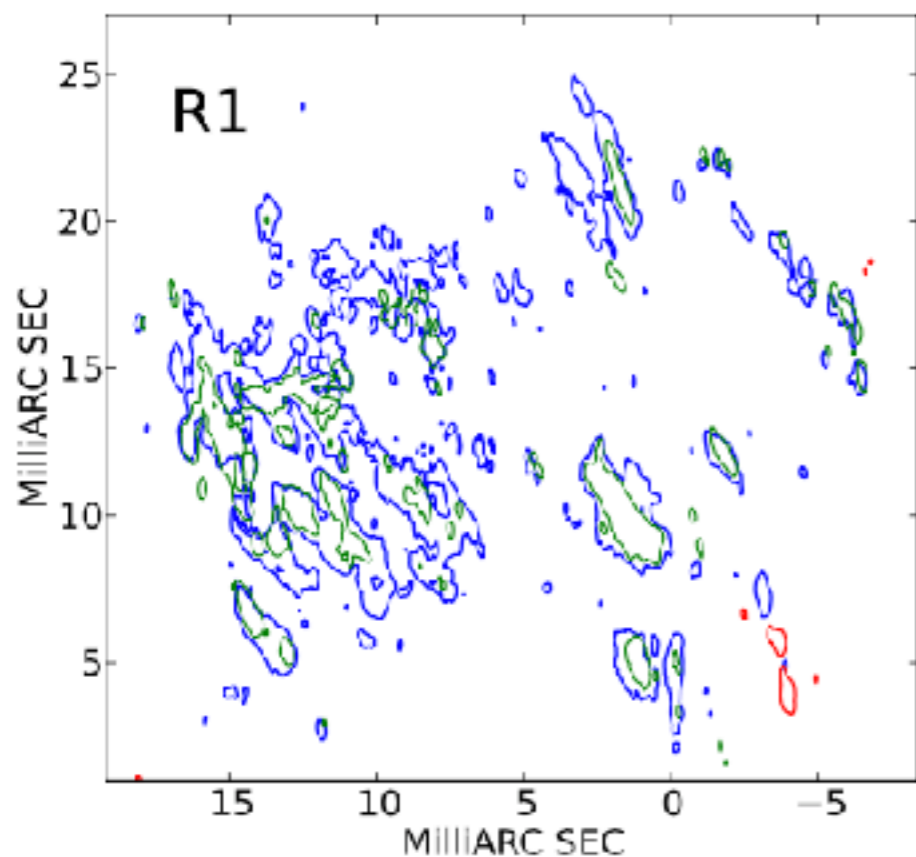
Very well overlapped $v=1$ and $v=2$ $J=1-0$
 -> predominant collisional pumping
 -> line overlap via radiative pumping

R1: very weak linear pol.
 $v=1$ four features polarized ($\sim 10\%$)
 $v=2$ no pol.



Blue: $v=1, J=1-0$
 Green: $v=2, J=1-0$
 Red: $v=1, J=2-1$

Results and discussion



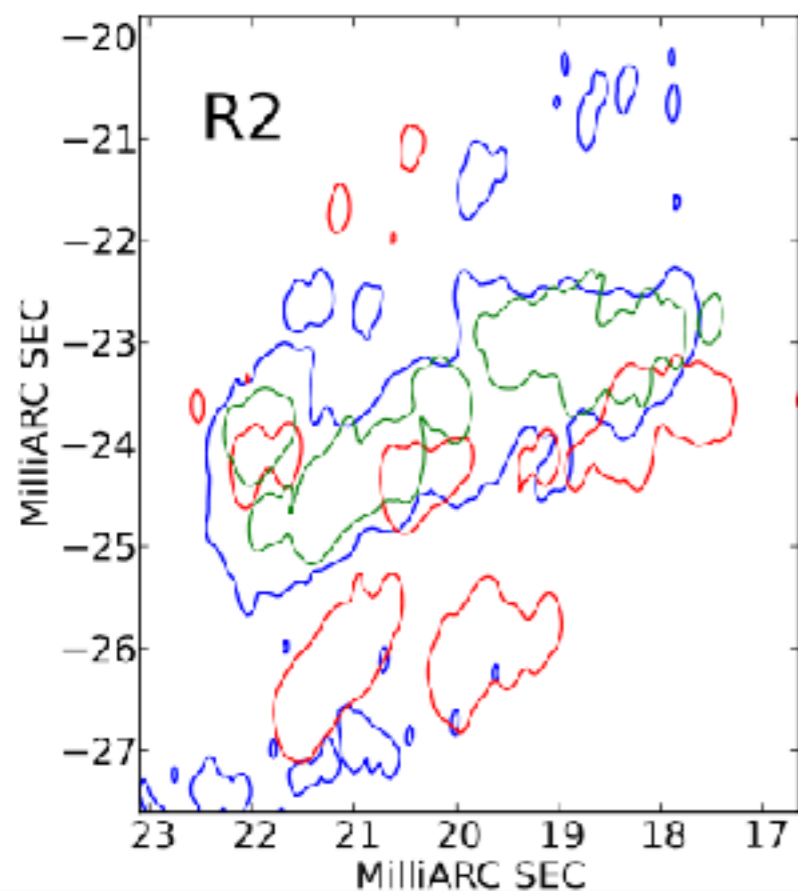
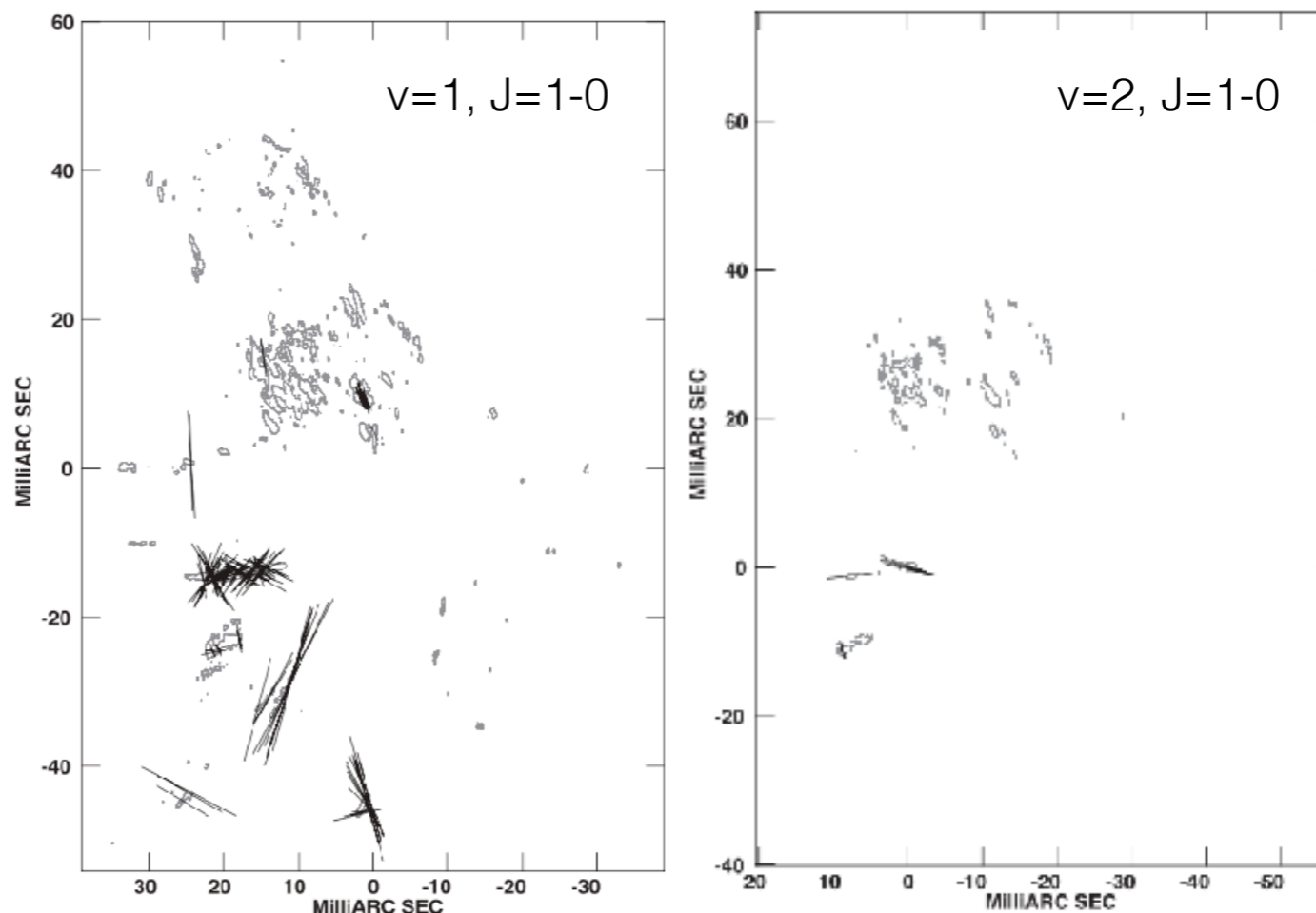
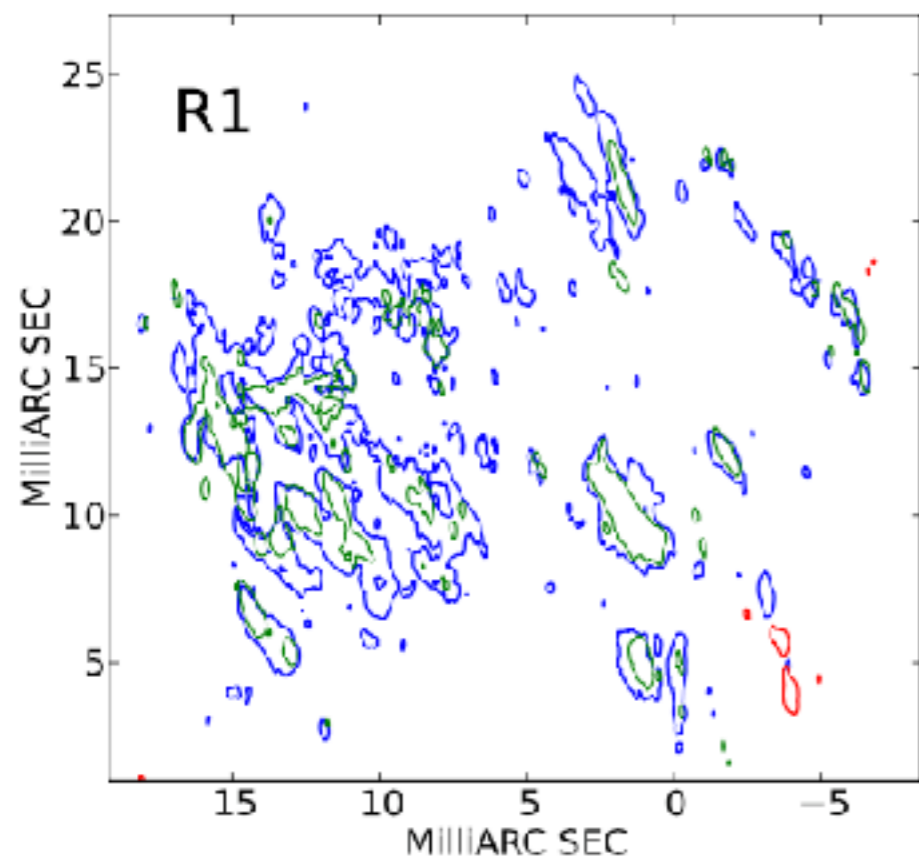
Linear pol.: natural characteristics of radiative pumping (being weaker further from central star)
Wide radial extent of the maser emission in R1 with weak linear pol. of $v=1$ and no pol. of $v=2$
-> contrary to radiative pumping scheme

Blue: $v=1, J=1-0$

Green: $v=2, J=1-0$

Red: $v=1, J=2-1$

Results and discussion



R2

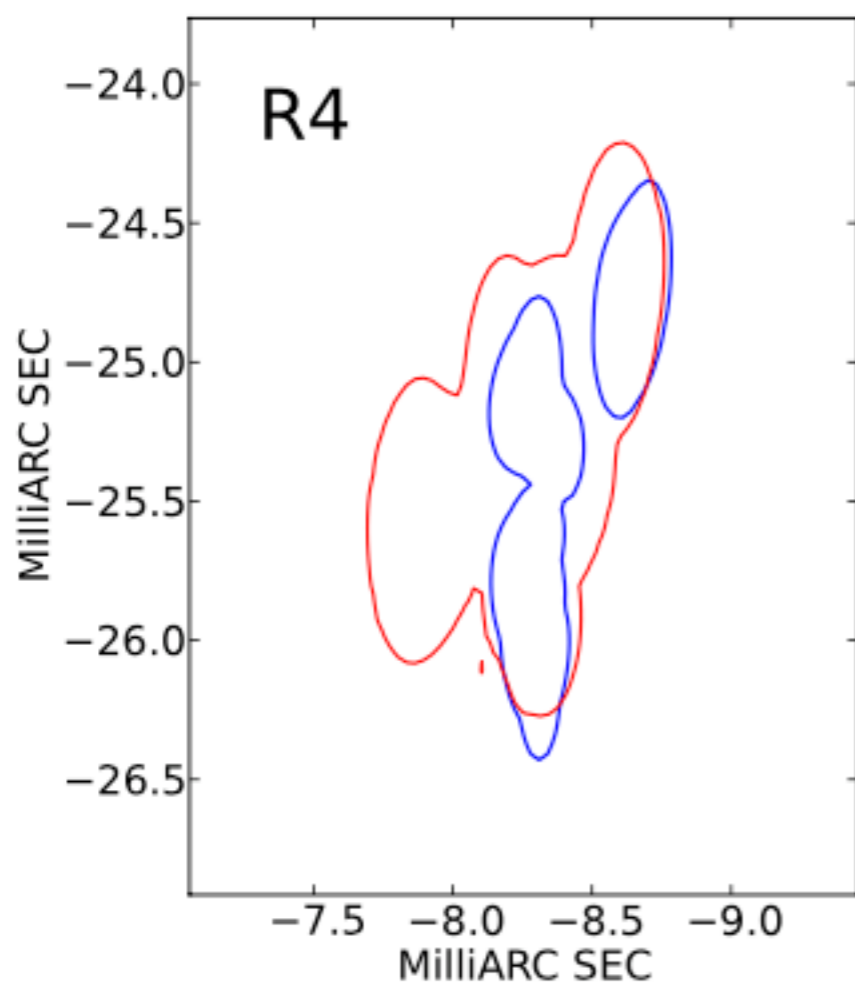
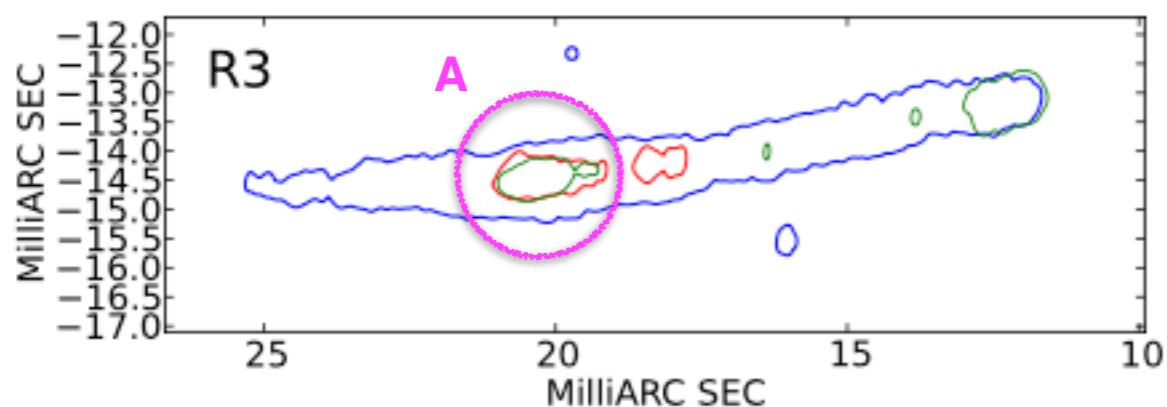
Very strong spatial coincidence of $v=1$ and $v=2$
IF line overlap mechanism causes this
THEN observed spatial overlap between
the $v=1$ $J=1-0$ and $v=1$ $J=2-1$ lines would only occur
under a very restricted set of envelope conditions.

Blue: $v=1, J=1-0$

Green: $v=2, J=1-0$

Red: $v=1, J=2-1$

Results and discussion



R3

The $v=2$ $J=1-0$ emission is located within two subsets of the $v=1$ $J=1-0$ emission regions

one: at the peak, the other: inner end

$v=2$ is closet to the star

-> rad. pumping: line overlap

-> coll. pumping: higher density condition

All transitions occur and are overlapped at A.

-> Longest coherent path (gain) length at A

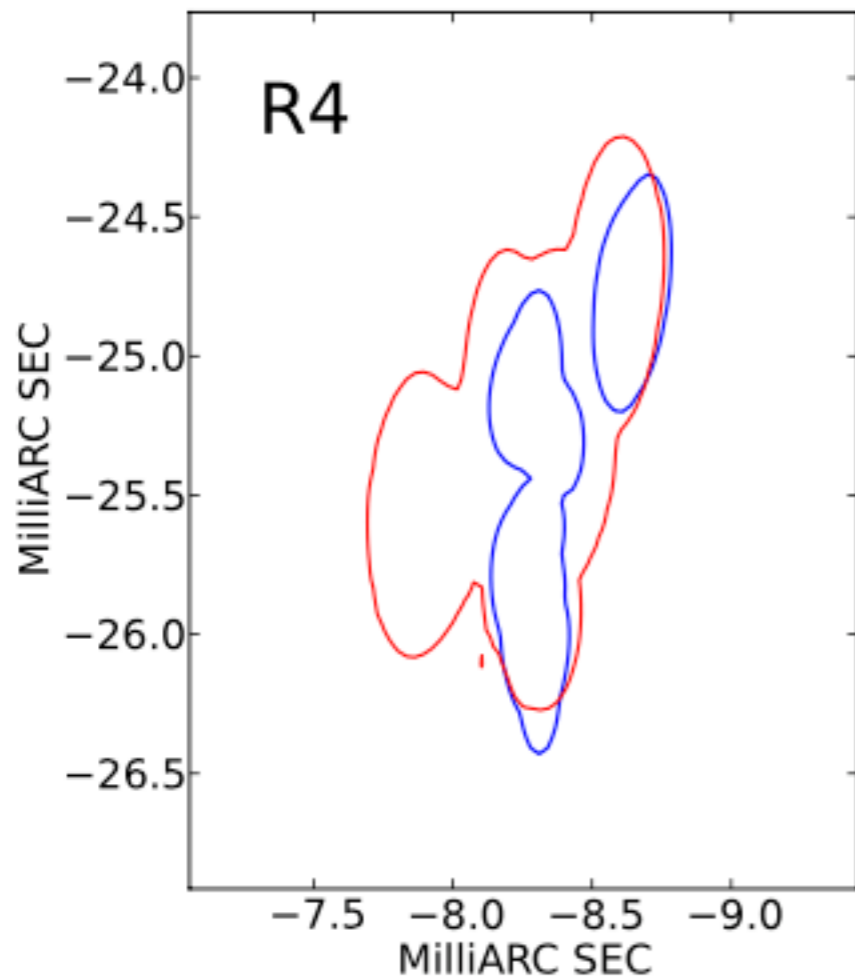
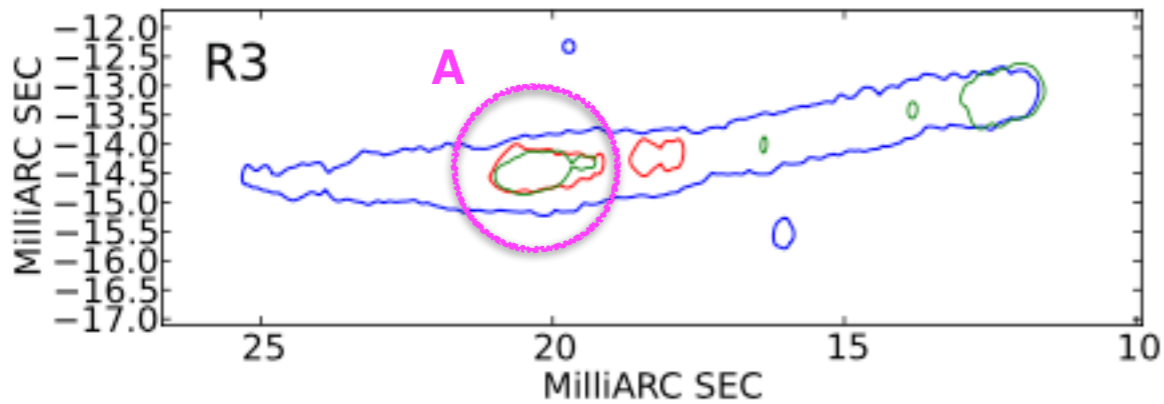
-> Competitive gain effect (maser saturation causes the coupling of the different vibrational transitions)

Blue: $v=1$, $J=1-0$

Green: $v=2$, $J=1-0$

Red: $v=1$, $J=2-1$

Results and discussion



R3

Nedoluha and Watson (1990): linear pol. decreasing with maser saturation

Western and Watson (1984): linear pol. increasing with maser saturation

Elitzur (1991): linear pol. level is the same in the unstarated and saturated masers

In this work, strong linear pol. of all transitions in R3

J=1-0: 1~20 %

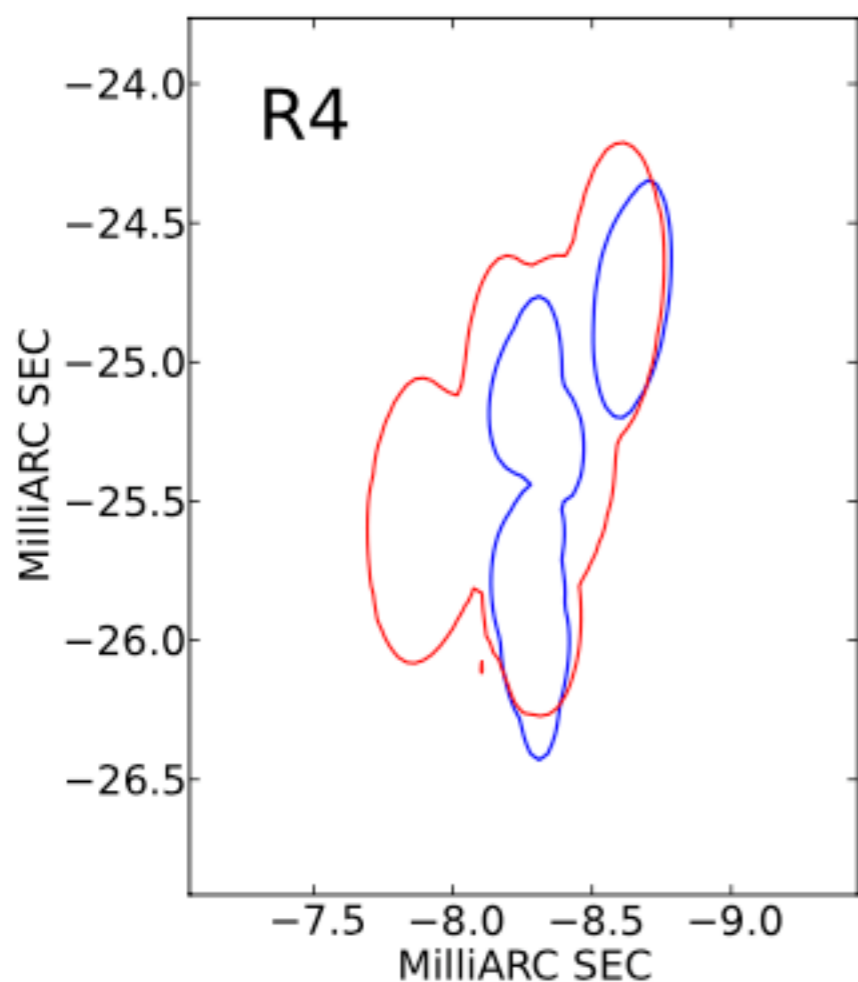
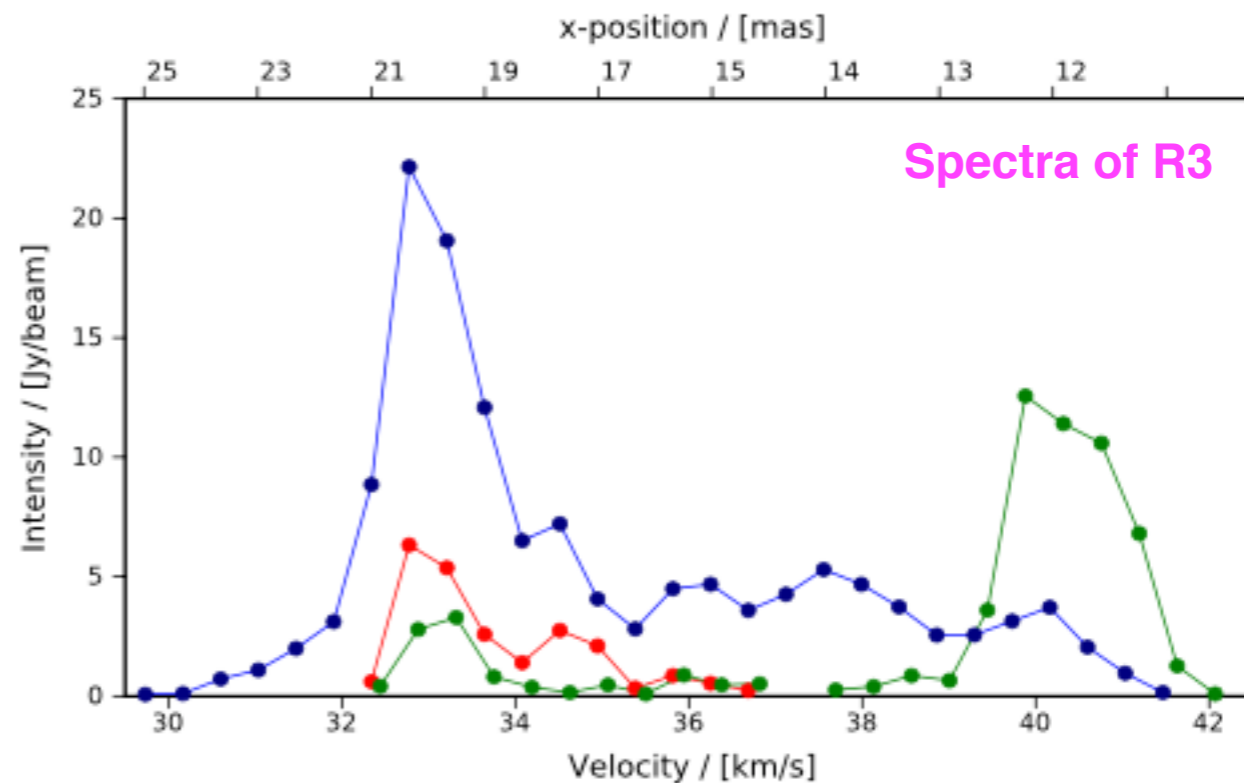
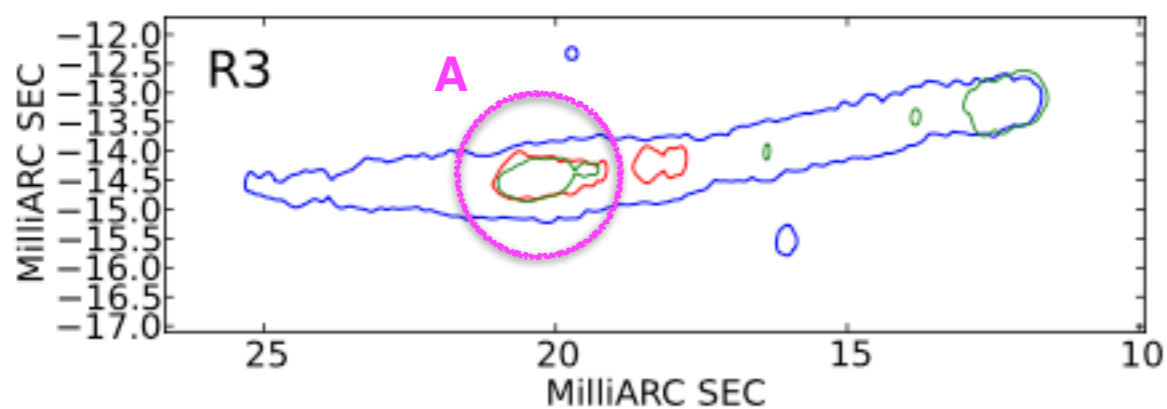
J=2-1: 25~40 %

Blue: $v=1, J=1-0$

Green: $v=2, J=1-0$

Red: $v=1, J=2-1$

Results and discussion



$v=1$ $J=1-0$ and $v=1$ $J=2-1$

Generally little overlap have been found.

In this work, several overlapped regions are found (in R2 and R3).

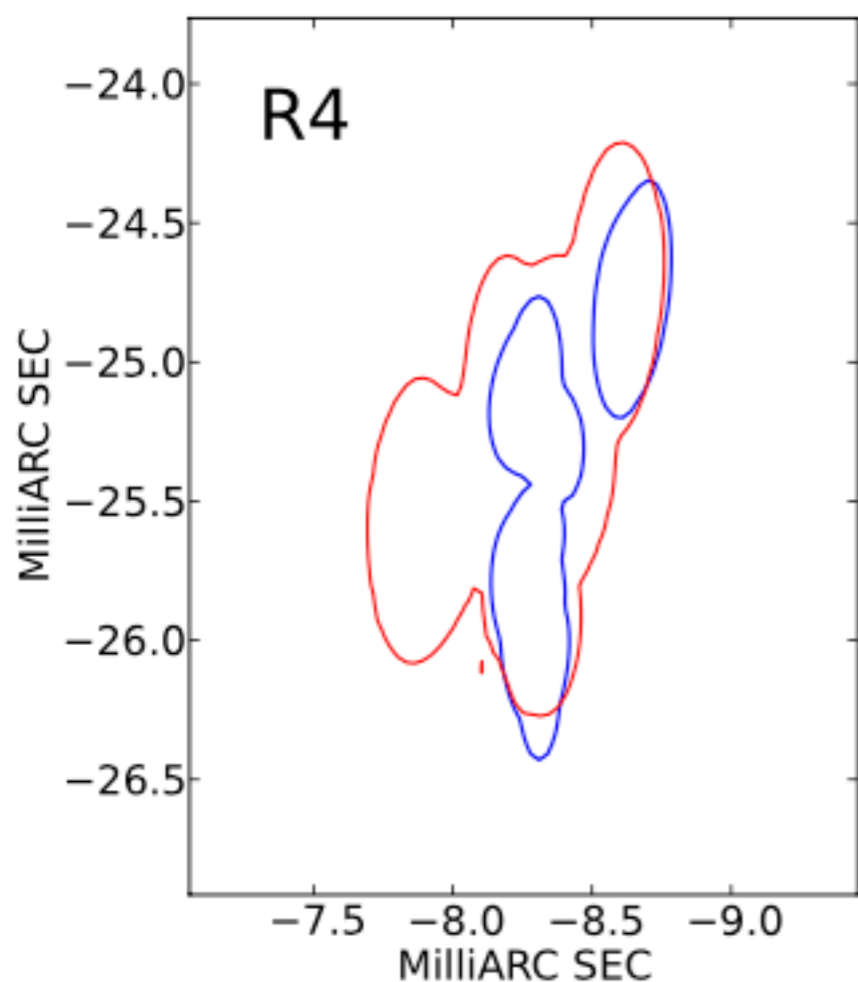
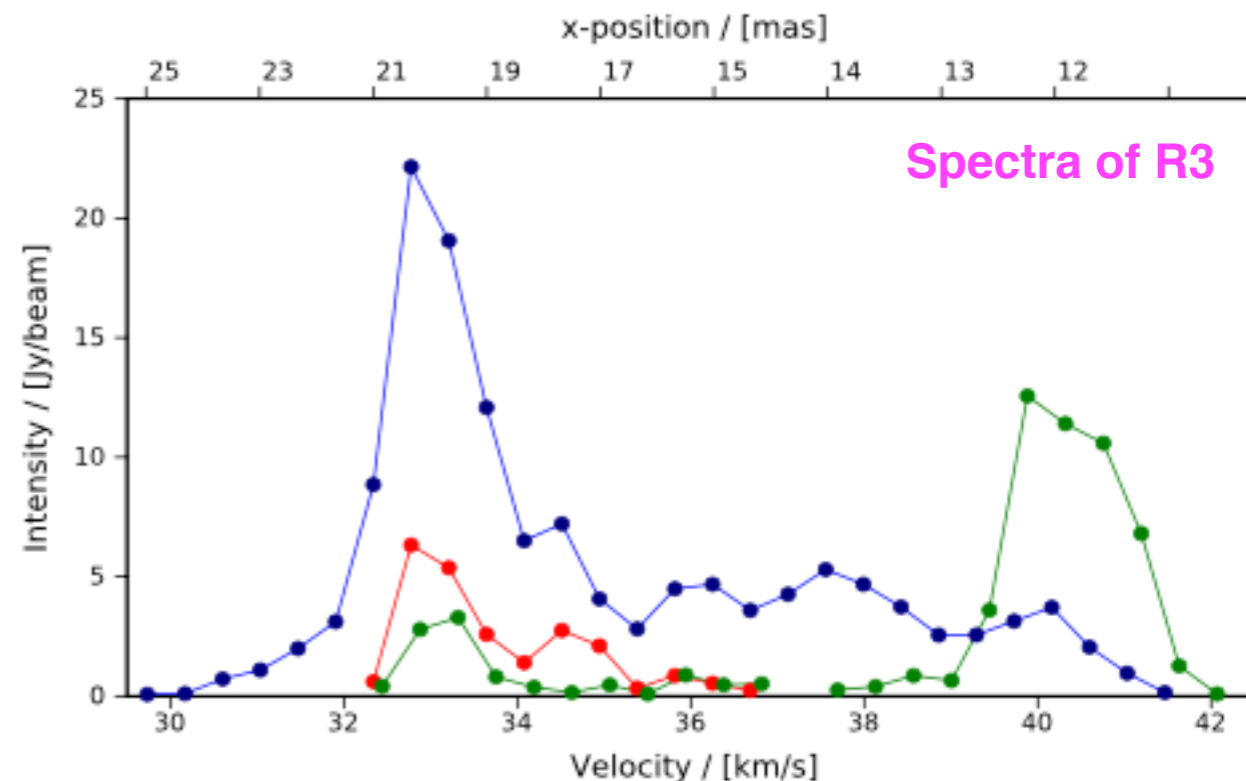
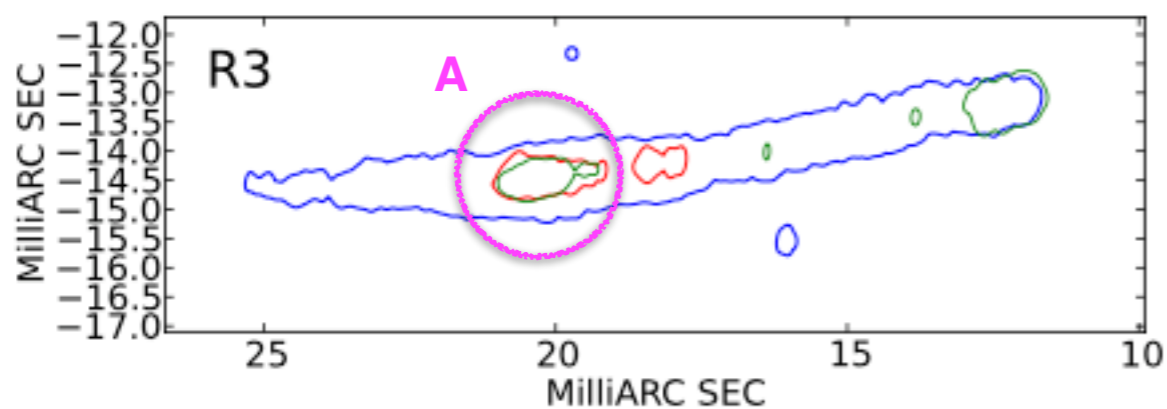
Similar spectra shapes of $v=1$ $J=1-0$ and $v=1$ $J=2-1$
-> expect similar spatial distribution
-> found the similar spatial overlap only in R3

Blue: $v=1$, $J=1-0$

Green: $v=2$, $J=1-0$

Red: $v=1$, $J=2-1$

Results and discussion



Alcolea (2004): coincidence of $v=1$ $J=1-0$ and $v=1$ $J=2-1$

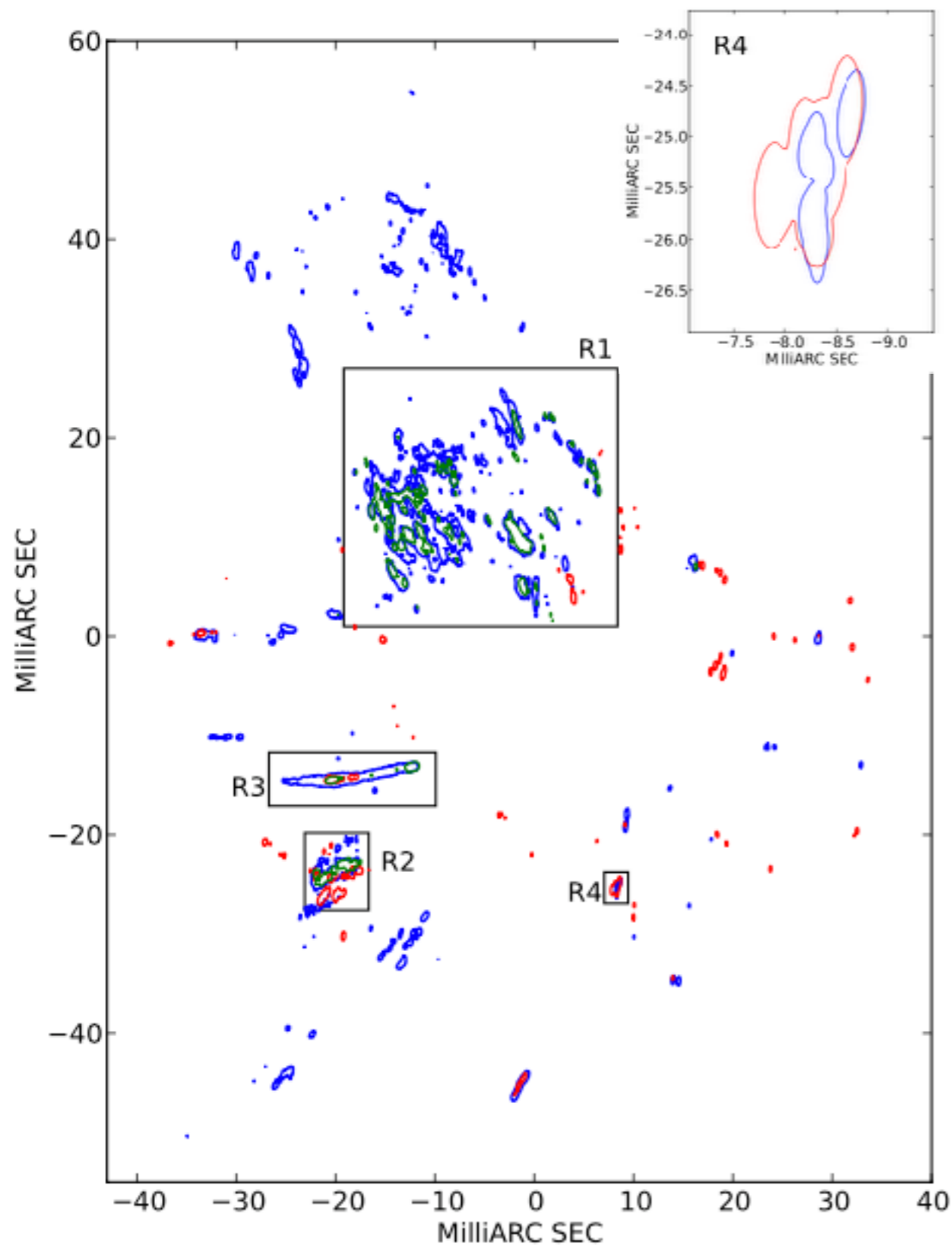
Humphreys (2002): stronger intensity of $J=2-1$ than $J=1-0$

Doel (1995): predominant coll. pumping
 $J=2-1$ is stronger than $J=1-0$ for intermediate density
 For high and low density: $J=2-1$ is weaker than $J=1-0$

Weaker $J=2-1$ of this work \rightarrow low density in overlap regions (consistent with Phillips 2003 for R Cas)

Blue: $v=1$, $J=1-0$
 Green: $v=2$, $J=1-0$
 Red: $v=1$, $J=2-1$

Results and discussion



Doel (1995): Most of J=2-1 do not have J=1-0 counter part.

- > western part of overlapped map
- > imply intermediate density

R1

Doel (1995): notable absence of J=2-1
-> low density

IF line overlap via rad. pumping causes the close correspondence of $v=1$ and $v=2$ THEN absence of J=2-1 imply high density. (Soria-Ruiz 2004)

R4

$I(J=2-1) > I(J=1-0)$

J=2-1 is more spatially extended

No $v=2$ J=1-0 emission

Significant pol.

J=1-0: ~8%

J=2-1: ~19%