#### Discussion for a pre-study on Extended KVN

- Science Working Group -

Friday Seminar 2017 June 23



#### Plan for the pre-study on science

» Study on the distinguished capability of an extended KVN w.r.t. that of the KVN

- Improved imaging sensitivity due to the increased B (number of baselines)
  - N=4 —> B=6 : 2 x DR (the Dynamic range of the current KVN)
  - N=5 -> B=10 : 3 x DR
- Uniqueness of the additional baselines
  - KVN's longest baseline length: 300km
  - VLBA's shortest one : 400km
  - VLA' longest one : 40km
  - eMERLIN's longest one: 200km (upto 24GHz)
  - unique baselines: 40-300km (unsampled coverage at >24GHz)



#### Plan for the pre-study on science

» Study on unique science case candidates only with the distinguished capability

- Improved imaging sensitivity : faint sources or structures
  - high-z AGN
  - large scale jet structures of nearby AGN
  - microquasars
  - faint maser features of late type stars or star-forming regions
- Increased shorter baselines : large angular-scale sources
  - Galactic compact radio sources (radio stars, SNe, protostellar disks etc)

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- extragalactic compact radio sources (SNe, radio galaxies etc.)
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### **Discussion** plan

- Science WG meetings
  - KVN Science group leaders (June 12): suggestion of science cases
  - KVN internal discussion (Friday seminar, today): discussion, WG formation
  - Science WG meetings (July-October)
- · Collection of interests from KaVA Science mailing list: A. Jung, Imai, Zhao
- 1st workshop
  - Radio telescope user's meeting (August 17-18):
  - invitation of researchers from East Asian regions based on the inertests and the WG activity
- KAS meeting
  - · October: presenting results of the pre-study on scientific justifications
- 2nd workshop
  - KaVA/EAVN joint sceince workshop (Nov. 21-23)?
  - A workshop in honor of Dr. Cho's retirement (end Nov)?
- White Paper for science cases
  - · Korean version to fit in the formal report of the pre-study
  - · English version to be published to e.g., JKAS



# Plan for the White paper (Science cases)

- Complete list of authors (end June)
  - volunteer based and some suggestions
- Discussion in WG meetings (7-10월)
- Contents:
  - Introduction: KVN and EKVN
  - Review of current KVN sciences
  - Expected performance of EKVN
  - Science cases of EKVN
  - Summary
  - try to divide it into several white papers for e.g., overview, science cases



## Scientific projects using current KVN (limitations and future plans)

- Scientific projects using current capability of KVN
  - fringe detection (high sensitivity)
    - MASK, MASK2 with EKVN (amplitude self-cal.)
    - MQ
    - · AGN (GPS, Young Radio Galaxies)
    - · SFR-44G-Methanol
    - 95G Class I Methanol Maser with EKVN (shorter baseline)
    - 85G Class II Methanol Maser with EKVN (shorter baseline)
  - imaging capability (full-track or snapshot)
    - iMOGABA
    - MQ (amplitude self-cal)
    - AGN (GPS, FR0, YRG) (EKVN with amplitude self-cal)
    - SgrA\* with EKVN
  - · multi-wavelength accurate flux measurement
    - iMOGABA
    - MASK
  - 4(3)-band multi-frequency simultaneous observations
    - AGB, AGN, SFR-OrionKL, MQ
  - high resolution
  - high frequency : MQ (histerisis curve, at higher frequencies of > 22GHz), SiO masers at 86/129GHz
  - · phase referencing
  - polarization : 86/129GHz (SNU), 44/95G Methanal maser with EKVN
  - amplitude self-calibration
  - wide bandwidth: radio flare in protostar (astrometry)



# Suggested science cases by KVN science group leaders

- Science cases
  - Using unique shorter baselines (40-300km, @ >= 43GHz )
    - AGB star continuum, extended H2O/SiO maser features
    - MASK2
    - 95/85G Methanal Maser
    - microquasars (43/86/129GHz)
    - AGN (GPS) with radio lobe/hot spots (43/86/129GHz)
  - Using longer baselines (>500km, 86/129GHz)
    - high spatial resolution ops of SiO maser features
    - MASK2, microquasars (86/129GH)
    - AGN (GPS) with radio lobe/hot spots (86/129GHz)
  - Using improved imaging sensitivity:
    - HCN maser (80G대역), SiO 86G maser (v>=2), 29SiO, 30SiO
  - Large radio telescope (D~40m?):



### Discussion

- any volunteer for science WG
  - A. Jung, D. Byun, JW Lee
  - G. Zhao, J. Hodgson, J. Algaba
  - Imai, A. Jung
- any suggestions for science cases
  - astrometric capability based on improved imaging capability with additional longer baselines
  - better to suggest more detailed key science project with EKVN (e.g., imaging of protostellar disk with ALMA
  - Large radio telescope (D~40m?):
    - · better to study in more detail
    - may need broader science cases with more careful discussion/efforts/pre-study
    - a thought on having the EKVN with same size telescopes (e.g. wide field imaging, etc.) :
    - many people want to have larger KVN telescope for e.g., extragalactic sciences (CO survey)
  - possibility to expand the backend of the KVN within this project?
    - aim to have this project a 5year project which may not fit to the expansion



### discussion

- possibility to have additional stations abroad
  - budget limitation? 30billion KRW
- need to distinguish the EKVN sciences with those of KaVA
- a long baseline, high cadence AGN cosmology
  - would be better to have multi-frequency capability

